

ON THE MOTION OF A BODY OF REVOLUTION BOUNDED BY A SPHERE, ON A SPHERICAL BASE

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This paper is closely related to [1 to 4]. The problem considered here is that of rolling of the body of revolution bounded by a sphere, on another fixed sphere, in case when the force function of the given forces is a function of u coordinate only. In other words, we shall investigate the equations*

$$\frac{d}{du} \frac{\partial \Theta}{\partial \sigma} + \frac{R}{R_c} \frac{\partial \Theta}{\partial n} - n [M(R^2 + l^2 + 2Rl \cos u) + A] \frac{R - R_c}{R_c} + \sigma [M(R^2 + l^2 + 2Rl \cos u) + A] \operatorname{ctg} u + MlR \sin u \sigma + \kappa \frac{R - R_c}{R_c} \cos u = 0 \quad (1)$$

$$\frac{d}{du} \frac{\partial \Theta}{\partial n} + \frac{R}{R_c} \sigma [M(R^2 + l^2 + 2Rl \cos u) + A] - \frac{R}{R_c} \frac{\partial \Theta}{\partial \sigma} + MlR \sin u + \kappa \frac{R - R_c}{R_c} \sin u = 0$$

integration of which is equivalent to the problem of determining the magnitudes u , ν , σ , π and n in the above problem.

Before anything else, we shall write the above equations in the more detailed form, assuming that the unknown functions are n and $r = -\sigma \sin u + n \cos u$ (projection of the angular velocity of the body on the Oz -axis), while $\gamma = \cos u$ is the independent variable. This yields the equalities

$$[M(R^2 + l^2 \gamma^2 + 2Rl\gamma) + A + (C - A)(1 - \gamma^2)] \frac{dr}{d\gamma} + [(C - A - Ml^2)\gamma - MlR] \frac{R - R_c r}{R_c} - [MlR\gamma^2 + (A + Ml^2 + MR^2)\gamma + MlR] \frac{dn}{d\gamma} - M(R + l\gamma) \frac{R^2}{R_c} n + \kappa \frac{R - R_c}{R_c} \gamma = 0 \quad (2)$$

$$[MlR - (C - A - Ml^2)\gamma] \frac{dr}{d\gamma} + (C - A - Ml^2) \frac{R - R_c}{R_c} r - (A + Ml^2 + MlR\gamma) \frac{dn}{d\gamma} - Ml \frac{R^2}{R_c} n + \kappa \frac{R - R_c}{R_c} = 0$$

* First and third equation of (2.5) in [3], where divided by u .

Multiplying the second equation by $-\cos u$ and adding it to the first one, we obtain

$$[C + MR(R + l\gamma)] \frac{dr}{d\gamma} - MR(l + R\gamma) \frac{dn}{d\gamma} - MlR \left(\frac{R}{R_c} - 1 \right) r - M \frac{R^3}{R_c} n = 0 \quad (3)$$

while from the second equation of (2) together with (3), we obtain

$$\frac{dn}{d\gamma} = \frac{1}{AR} [R(A - C)\gamma - lC] \frac{dr}{d\gamma} + \frac{C - A}{A} \left(\frac{R}{R_c} - 1 \right) r + \frac{\kappa}{A} \left(\frac{R}{R_c} - 1 \right) \quad (4)$$

$$\frac{d^2n}{d\gamma^2} = \frac{1}{AR} [R(A - C)\gamma - lC] \frac{d^2r}{d\gamma^2} + \frac{C - A}{A} \left(\frac{R}{R_c} - 2 \right) \frac{dr}{d\gamma}$$

Differentiating (3) with respect to γ and inserting (4), we find the following non-homogeneous linear second order equation for r

$$[AC + MR^2A(1 - \gamma^2) + MC(l^2 + 2lR\gamma + R^2\gamma^2)] \frac{d^2r}{d\gamma^2} + 3[CMR(l + R\gamma) - AMR^2\gamma] \frac{dr}{d\gamma} - MR^2(C - A) \left(\frac{R^2}{R_c^2} - 1 \right) r - MR^2\kappa \left(\frac{R^2}{R_c^2} - 1 \right) = 0 \quad (5)$$

Corresponding homogeneous equation can be written in the form (a_1, a_2, a_3, b_1, b_2 and c_1 are constants)

$$(a_1x^2 + a_2x + a_3) \frac{d^2r}{dx^2} + (b_1x + b_2) \frac{dr}{dx} + c_1r = 0 \quad (6)$$

and is of Fuchsian type [5]. Its singularities are $c' = \infty$ and the roots a' and b' of the equation

$$a_1x^2 + a_2x + a_3 = 0 \quad (7)$$

Assuming that $r = z(y)$ and $x = a' + (b' - a')\gamma$ (in case $a' \neq b'$), we obtain a hypergeometric equation

$$y(1 - y) \frac{d^2z}{dy^2} + \left(\frac{3}{2} - 3y \right) \frac{dz}{dy} - \left(1 - \frac{R^2}{R_c^2} \right) z = 0 \quad (8)$$

Now, first equation of (4) together with the relation $r = -\sigma \sin u + n \cos u$, enables us to find n and σ in terms of u . These can then be inserted into the force integral. This reduces the problem of determination of u as a function of time, to a quadrature, and consequently settles the problem of determination of u, v, σ, τ and n .

In this manner the problem of rolling of a body of revolution bounded by a sphere on another fixed sphere when the given force function depends only on u , is reduced to integration of one Riccati equation and to quadratures [2 and 3].

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